High redshift deuterium abundance

Resolving the tension between quasar and CMB Ω_b measurements

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Outline

Motivation

Voigt Function

Kramers-Heisenberg Formula

Voigt & K-H

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Table 4. The sample of D1/H1 measurements considered robust in Pettini et al. (2008) together with updated estimates in the same absorbers and more recent, similarly precise measurements from other absorbers.

| Reference | Absorption redshift | $\log(N(HI))$ | [X/H] | D I/H I [×10 ⁻⁵] | $100\Omega_b h^2$ |
|---|------------------------|--------------------|----------|---------------------------------|-------------------|
| Burles & Tytler (1998a) | 2.504 | 17.4 ± 0.07 | -2.55 Si | 4.00 ± 0.70 | 1.66 ± 0.18 |
| Pettini & Bowen (2001) | 2.076 | 20.4 ± 0.15 | -2.23 Si | 1.65 ± 0.35 | 2.82 ± 0.36 |
| Kirkman et al. (2003) | 2.426 | 19.7 ± 0.04 | -2.79 O | 2.43 ± 0.35 | 2.24 ± 0.20 |
| Fumagalli et al. (2011) | 3.411 | 18.0 ± 0.05 | -4.20 Si | 2.04 ± 0.61 | 2.49 ± 0.05 |
| Noterdaeme et al. (2012) | 2.621 | 20.5 ± 0.10 | -1.99 O | 2.80 ± 0.80 | 2.05 ± 0.35 |
| Cooke et al. (2014), Pettini & Cooke (2012) | 3.050 | 20.392 ± 0.003 | -1.92 O | 2.51 ± 0.05 | 2.19 ± 0.02 |
| Cooke et al. (2014), O'Meara et al. (2001) | 2.537 | 19.4 ± 0.01 | -1.77 O | 2.58 ± 0.15 | 2.16 ± 0.04 |
| Cooke et al. (2014), Pettini et al. (2008) | 2.618 | 20.3 ± 0.01 | -2.40 O | 2.53 ± 0.10 | 2.18 ± 0.03 |
| Cooke et al. (2014) | 3.067 | 20.5 ± 0.01 | -2.33 O | 2.58 ± 0.07 | 2.16 ± 0.03 |
| Cooke et al. (2014), O'Meara et al. (2006) | 2.702 | 20.7 ± 0.05 | -1.55 O | 2.40 ± 0.14 | 2.25 ± 0.03 |
| Riemer-Sørensen et al. (2015) | 3.255 | 18.1 ± 0.03 | -1.87 O | 2.45 ± 0.28 | 2.23 ± 0.16 |
| Balashev et al. (2016) | 2.437 | 19.98 ± 0.01 | -2.04 O | 1.97 ± 0.33 | 2.54 ± 0.26 |
| This work | 3.572 | 17.925 ± 0.006 | -2.26 O | 2.62 ± 0.05 | 2.14 ± 0.03 |
| Weighted average ¹ | | | | 2.55 ± 0.03 | 2.17 ± 0.03 |
| Unweighted average ¹ | _ | | _ | 2.53 ± 0.17 | 2.18 ± 0.08 |
| Planck Collaboration et al. (2016) | _ | _ | _ | 2.45 ± 0.05 | 2.225 ± 0.016 |

The conversion between D1/H1 and $\Omega_b h^2$ is based on nuclear rates from Coc et al. (2015) for standard Big Bang Nucleosynthesis. ¹Without the Balashev et al. (2016) and Noterdaeme et al. (2012) measurements

Figure: Riemer-Sørensen et al. (MNRAS 468, 3239)

D/H measurment

Lyman Series:

$$E_n = \frac{\mu e^4}{2(4\pi\epsilon_0\hbar)^2} \frac{1}{n^2}$$
 and $\mu = \frac{Mm_e}{M+m_e}$,

where n is the principle quantum number, m_e and M are the mass of electron and nucleus.

- \Box Hydrogen and Deuterium: $\mu_H = \frac{1836}{1+1836}$ and $\mu_D = \frac{3670}{1+3670}$
- \Box Lyman $\alpha:~\lambda_H=1215.67$ Å, $\lambda_D=1215.34$ Å

D/H measurment

- Lyman α : $\lambda_H = 1215.67$ Å, $\lambda_D = 1215.34$ Å
- Left: $n_H = 10^{16}/cm^2$, $n_D/n_H = 10^{-3}$ (solid), 10^{-4} (dashed)
- Right: $n_H = 10^{16}$ (dotted), 10^{18} (dashed) and 10^{19} (solid) with $n_D/n_H = 10^{-4}$



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Cross-section of a two-level system:

$$\sigma(\nu) = cf \sqrt{\frac{3\pi\sigma_T}{8}} \phi(\nu) \,,$$

where ν is the incoming photon energy, f is the oscillator strength, $\phi(\nu)$ is the Lorentzian

$$\phi(\nu) = \frac{\Gamma/4\pi^2}{(\nu - \nu_0)^2 + (\Gamma/4\pi)^2},$$

 $E = h\nu_0$, and Γ is the spontaneous decay rate.

- Doppler effect: $\nu_0 \rightarrow \nu_0 \left(1 + \frac{v_z}{c}\right)$
- Boltzmann distribution: $N(v_z) = \frac{exp(-v_z^2/(2kT/m))}{\sqrt{2\pi kT/m}}$



 \blacksquare Intensity Profile: $I_{\nu}=I_{0}e^{-N\alpha_{\nu}}$, and

$$\alpha_{\nu} = \frac{\sqrt{\pi}e^2}{m_e c} \frac{f}{\Delta\nu_D} H(a, u) \,,$$

where
$$\Delta \nu_D = \frac{b\nu_0}{c}$$
, $b = \sqrt{\frac{2kT}{m}}$, $u = \frac{(\nu - \nu_0)}{\Delta \nu_D}$ and $a = \frac{\Gamma}{4\pi \Delta \nu_D}$.
Voigt Function:

$$H(a, u) = \frac{a}{\pi} \int_{-\infty}^{\infty} \frac{e^{-y^2} dy}{(u-y)^2 + a^2} \,.$$

- Absorption and Voigt Function:
 - $\hfill\square$ Two-level system \Rightarrow Lorentzian profile
 - Convolution of Boltzmann and Lorentzian

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Kramers-Heisenberg formula

$$\frac{d\sigma}{d\Omega} = r_0^2 \left(\frac{\nu'}{\nu}\right) \left| \delta_{if} \bar{\epsilon}^{\alpha} \bar{\epsilon}^{\alpha'} + \frac{2\pi m_e \nu_{nf} \nu_{ni}}{\hbar} \right. \\ \left. \sum_n \left[\frac{(\vec{x} \cdot \bar{\epsilon}^{\alpha'})_{fn} (\vec{x} \cdot \bar{\epsilon}^{\alpha})_{ni}}{\nu_{ni} - \nu - i\Gamma_n/2} + \frac{(\vec{x} \cdot \bar{\epsilon}^{\alpha})_{fn} (\vec{x} \cdot \bar{\epsilon}^{\alpha'})_{ni}}{\nu_{ni} + \nu'} \right] \right|^2,$$

where $\nu(\epsilon^\alpha)$ and $\nu'(\epsilon^{\alpha'})$ are the frequency (orientation) of the incoming and outgoing photon, and

$$\nu_{ab} = \frac{E_a - E_b}{h} \,.$$

• Lyman α : $1s \rightarrow 2p \rightarrow 1s$



Figure: Kiehunn Bach, Hee-Won Lee, JKAS 47, no.5, 187(2014)

Kramers-Heisenberg formula

$$\frac{d\sigma}{d\Omega} = r_0^2 \left(\frac{\nu'}{\nu}\right) \left| \delta_{if} \bar{\epsilon}^{\alpha} \bar{\epsilon}^{\alpha'} + \frac{2\pi m_e \nu_{nf} \nu_{ni}}{\hbar} \right| \\ \sum_n \left[\frac{(\vec{x} \cdot \bar{\epsilon}^{\alpha'})_{fn} (\vec{x} \cdot \bar{\epsilon}^{\alpha})_{ni}}{\nu_{ni} - \nu - i\Gamma_n/2} + \frac{(\vec{x} \cdot \bar{\epsilon}^{\alpha})_{fn} (\vec{x} \cdot \bar{\epsilon}^{\alpha'})_{ni}}{\nu_{ni} + \nu'} \right] \right|^2.$$

- Rayleigh Scattering (Elastic): $1s \rightarrow np \rightarrow 1s$
- Raman Scattering (Inelastic): $1s \rightarrow np \rightarrow n's$ or n'd
- Example: $1s \rightarrow 4p \rightarrow$ Final states (1s, 2s, 3s and 3d)

•
$$\sigma(\nu) = \sigma^{Rayleigh}(\nu) + \sum_{f} \sigma_{f}^{Raman}(\nu).$$

• Doppler effect:
$$\nu_{ni} \rightarrow \nu_{ni} \left(1 + \frac{v_z}{c}\right)$$
.

Boltzmann distribution:
$$N(v_z) = \frac{exp(-v_z^2/(2kT/m))}{\sqrt{2\pi kT/m}}$$
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Lyman Series:



• Lyman α :



Lyman β:



• Lyman γ :



• Lyman α with $\log N = 13.0$:



• Lyman α with $\log N = 21.6$:



• Lyman β with $\log N = 21.6$:



• Lyman γ with $\log N = 21.6$:



Fitting

•
$$\delta(D/H) \equiv \frac{n_{voigt-fit}}{n_{fid}} - 1$$

- Fitting Region:
 - $\hfill\square$ blue: Ly α Ly14
 - $\hfill\square$ red: Ly α Ly12
 - $\hfill\square$ yellow: Ly α Ly7



Fitting

• Fit the fiducial with Ly α - Ly7

 $\hfill\square$ Left: Lyman 7 ($n=1\rightarrow 8$) with $\log N=20.6$

 \square Right: Lyman 7 ($n = 1 \rightarrow 8$) with $\log N = 21.6$



Summary

Deuterium Abundance

- □ QSO: $D/H = (2.55 \pm 0.03) \times 10^{-5}$
- □ CMB: $D/H = (2.45 \pm 0.05) \times 10^{-5}$

Speed:

- □ Voigt Profile: quick!!
- □ Karmers-Heisenberg: slow!!